Lecture 2

Cosmological Redshift and Galaxies

Galaxies

- The basic structures of the Universe at large scales.
- Galaxies are "cities" of stars.
- Our Milky Way Galaxy:
 - 1,000 billion solar masses, 400 billion stars, 100,000 lyrs radius





Edge-on: a thin disk

Face-on: a grand spiral

Types of Galaxies







Spiral Galaxy young & old stars Elliptical Galaxy old stars Irregular Galaxy a mixture

Andromeda





Our "Local Group" of galaxies Dominated by Milky Way & Andromeda



Virgo cluster ~2,000 galaxies, 16 Mpc away We are falling towards here at 750 km/s!

Cosmological Redshift: I



For distant galaxies we find that spectral lines are shifted to significantly longer wavelengths.

Cosmological Redshift: II

 As before, we define the Cosmological Redshift parameter, z, as the fractional change in wavelength

$$z = \frac{\Delta \lambda}{\lambda}$$

• However, for <u>distant galaxies</u> we find that

Space is expanding between us and the distant galaxy

- i.e. the distance between Earth and the distant galaxy is increasing, <u>not</u> that the galaxy is moving away from us!
- This comes from Hubble's Law see the next lecture.
 - See also today's tutorial

Cosmological Redshift: III

- When the velocity of expansion of the Universe is small
 - i.e. *v* << *c*, or *v*/*c* << 1, then

$$z = \frac{\Delta \lambda}{\lambda} \approx \frac{v}{c}$$

• The <u>very same formula</u> as for the Doppler Redshift, but now the interpretation is that <u>space is being stretched</u> at a speed *v*

• Note: there is an exact formula, but you don't need to know for A-level:

$$1 + z = \gamma (1 + \frac{v}{c})$$
 where $\gamma = 1/\sqrt{1 - \left(\frac{v}{c}\right)^2}$

• When $v/c \ll 1$ this reduces to the formula above, z=v/c.

Cosmological Redshift: IV

- Doppler Redshift applies for:
 - Motions of planets in our Solar System
 - Motions of stars in our Galaxy
 - Motions of nearby galaxies
- Cosmological Redshift applies for:
 - The stretching of space to distant galaxies